

# Towards new relativistic hydrodynamics from AdS/CFT

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with Edward Shuryak

- QGP is Deconfined
- QGP is strongly coupled (sQGP)  
behaves “almost” like a perfect liquid (Navier-Stokes with very small viscosity)  
 $\eta \sim \text{mean free path} \sim 1/\sigma$

QCD  $\longrightarrow \mathcal{N} = 4$  SYM (CFT)

Strong coupling (and large  $N_c$ )  $\rightarrow$  AdS/CFT  $\rightarrow$  SUGRA on AdS<sub>5</sub>

CFT at finite Temperature  $\leftrightarrow$  AdS Black Hole

## Outline of the talk

- Quick tour to Relativistic hydro
- Brief visit into 5th dimension: Black Hole AdS/CFT
- All order hydro: momenta dependent viscosity

**Motivation:** Experiments (RHIC) probe systems with finite gradients.

**Main Goal:**

Introduce higher order dissipative terms in the gradient expansion of  $T^{\mu\nu}$

Extract momenta dependent viscosities by matching two-point correlation functions of stress energy tensor with correlation functions computed from BH AdS/CFT.

(when applying to QCD we hope for some universality for transport coefficients)

We propose to use this hydro as a “nonlinear model” for real simulations at RHIC

# Relativistic Hydrodynamics

from Landau & Lifshitz V6

## Energy momentum tensor

$$T^{\mu\nu} = (\epsilon + P) u^\mu u^\nu + P g^{\mu\nu} + \Pi^{\mu\nu}$$

$u$  - velocity field of the fluid  $u^2 = -1$   $P$  - Pressure

$\Pi^{\mu\nu}$  - tensor of dissipations ( ideal fluid:  $\Pi^{\mu\nu} = 0$  )

$u_\mu \Pi^{\mu\nu} = 0$  - no dissipation in the local rest frame

## Navier Stokes term (expanding in the velocity gradient)

$$\Pi^{\mu\nu} = -\eta (\Delta^{\mu\lambda} \nabla_\lambda u^\nu + \Delta^{\nu\lambda} \nabla_\lambda u^\mu - \frac{2}{3} \Delta^{\mu\nu} \nabla_\lambda u^\lambda) - \xi \Delta^{\mu\nu} \nabla_\lambda u^\lambda$$

$$\Delta^{\mu\nu} = g^{\mu\nu} + u^\mu u^\nu$$

## Energy - momentum conservation:

$$\nabla_{\mu} T^{\mu\nu} = 0 \quad \longrightarrow \quad \text{Navier - Stokes Eq.}$$

## Conformal invariance

$$T^{\mu}_{\mu} = 0 \quad \longrightarrow \quad \epsilon = 3 P \quad \text{and} \quad \xi = 0$$

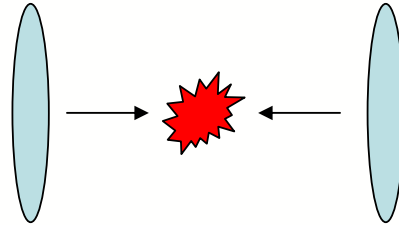
## Entropy density and EoS

$$s = \frac{\epsilon + P}{T} = 4 k_{SB} T^3$$

## No dissipation no entropy production:

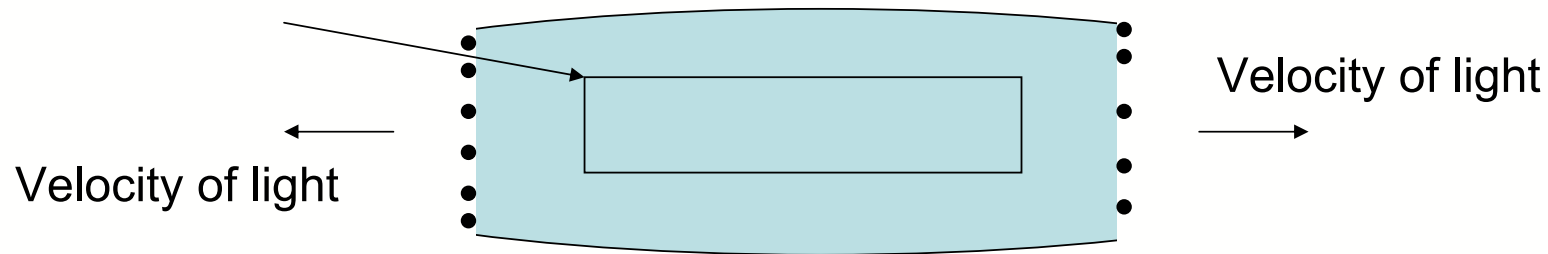
$$\frac{ds}{dt} = 0 \quad \text{if} \quad \Pi^{\mu\nu} = 0$$

# RHIC and Bjorken set up



Relativistically accelerated heavy nuclei

## Central Rapidity Region (CRR)



After collision

- **one-dimensional expansion: Boost Invariant Formulation**

**Local rest frame**  $u = (1, 0, 0, 0)$

$$(x^0, x^1, x_\perp) \rightarrow (\tau, y, x_\perp)$$

$\tau$  - proper time,  $y$  - spacetime rapidity

$$x^0 = \tau \cosh(y)$$

$$x^1 = \tau \sinh(y)$$

**The metric (1d Hubble expansion)**

$$ds^2 = -d\tau^2 + \tau^2 dy^2 + dx_\perp^2$$

**Hydro eq. simplify dramatically:**

$$\partial_\tau \epsilon(\tau) = -\frac{4\epsilon}{3\tau} + \frac{4\eta}{3\tau^2}$$

**Solution for  $\eta = 0$ :**

**Bjorken (1986)**

$$\epsilon \sim \frac{1}{\tau^{4/3}}$$

$$T \sim \frac{1}{\tau^{1/3}}$$

$$\partial_\tau (s\tau) = 0$$

# Sound waves in Relativistic Hydrodynamics

still from Landau & Lifshitz V6

Plane wave perturbation:

$$\delta u = \delta u_0 e^{-i\omega t + i q x}$$

$$\delta P = \delta P_0 e^{-i\omega t + i q x}$$

Linearized Hydro leads to the dispersion relation

$$\omega = c q - i \frac{2\eta}{sT} q^2$$

Sound velocity  $c = 1/\sqrt{3}$

Sound attenuation  $\sim \eta$

Spectral functions in the sound and shear channels ( $2\pi T = 1$  and  $\bar{\eta} \equiv 4\pi\eta/s$ )

$$\chi^L = \frac{2\omega c^2 q^4 \bar{\eta}}{(\omega^2 - c^2 q^2)^2 + 4\omega^2 c^2 q^4 \bar{\eta}^2}$$

$$\chi^T = \frac{\omega q^2 \bar{\eta}/2}{\omega^2 + q^2 \bar{\eta}^2/4}$$



# Israel-Stewart second order Hydrodynamics

Solves causality problems encoded in Navier-Stokes

Add extra term in the gradient expansion + non-linear terms in  $(\nabla u)$

$$\Pi^{\mu\nu} = (1 - \tau_R u_\lambda \nabla^\lambda) \Pi_{NS}^{\mu\nu}$$

Iterate the equation

$$(1 + \tau_R u_\lambda \nabla^\lambda) \Pi^{\mu\nu} = \Pi_{NS}^{\mu\nu}$$

When thinking about small perturbations  $u_\lambda \nabla^\lambda \rightarrow \nabla_t \rightarrow -i\omega$

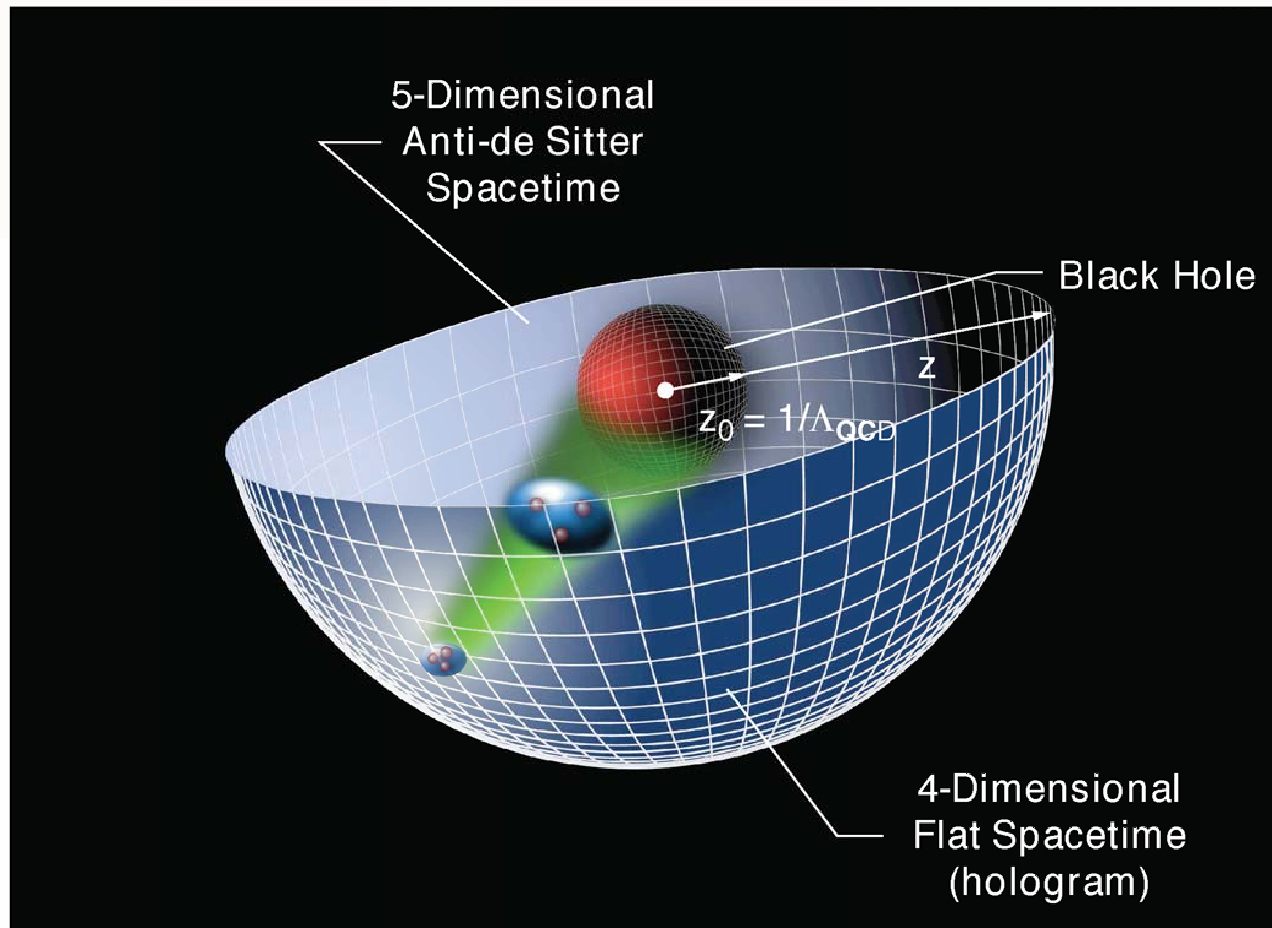
The IS second order hydro is equivalent (in the linear approximation) to

$$\eta \rightarrow \frac{\eta}{1 - i\tau_R \omega}$$

Sound dispersion

$$\omega = c q [1 + \bar{\eta} c^2 q^2 (2\tau_R - \bar{\eta})] - i c^2 \bar{\eta} q^2 [1 + c^2 q^2 \bar{\eta} \tau_R (2\bar{\eta} - \tau_R)]$$

# $AdS/QCD$



*Changes in  
length scale  
mapped to  
evolution in the  
5th dimension  $z$*

**AdS/CFT correspondence:** weakly coupled super-gravity in  $AdS_5 \times S_5$  is “dual” to strongly coupled  $\mathcal{N} = 4$  SYM gauge theory in 4d

**$AdS_5$  Schwarzschild BH metric**

$$ds^2 = \frac{\rho^2}{L^2} \left[ - \left( 1 - \frac{\rho_0^4}{\rho^4} \right) dt^2 + dx^2 + dy^2 + dz^2 \right] + \frac{L^2}{\rho^2 (1 - \rho_0^4/\rho^4)} d\rho^2$$

**BH Horizon at**  $\rho = \rho_0$

**AdS “boundary”**  $\rho \rightarrow \infty$  is Minkowski space (t,x,y,z)

**Gauge theory at the boundary is  $\mathcal{N} = 4$  SYM static plasma at finite temperature.**

**The Hawking temperature is**

$$T = \frac{\rho_0}{\pi L^2}$$

# Retarded correlators and Viscosity from AdS BH

Retarded correlators:

$$G_R^{\mu\nu\mu'\nu'}(\omega, q) = -i \int_0^\infty dt \int dx e^{-i\omega t + i q x} \langle [T^{\mu\nu}(t, x), T^{\mu'\nu'}(0, 0)] \rangle$$

**AdS/CFT:** energy-momentum tensor  $T_{\mu\nu}$  couples at the boundary to metric perturbations (gravitons). Solve linearized GR in 5d with absorptive boundary conditions at the horizon.

Shear viscosity

$$\eta = \lim_{\omega \rightarrow 0} \frac{1}{2\omega} \int dt dx e^{-i\omega t} \langle [T^{xy}(t, x), T^{xy}(0, 0)] \rangle$$

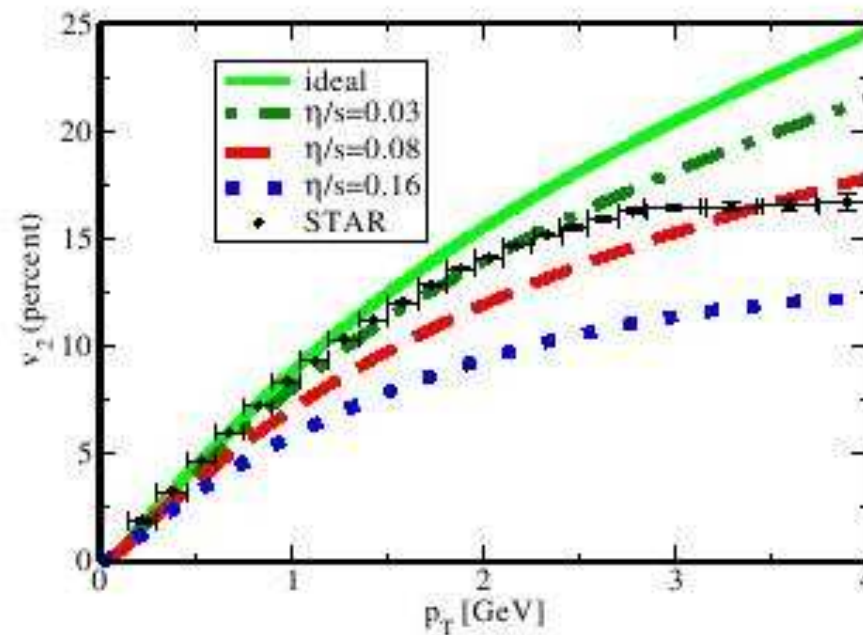
P. Kovtun, D. T. Son and A. O. Starinets, PRL 94, 111601 (2005)

$$\frac{\eta}{s} = \frac{1}{4\pi}$$

Imposing also Dirichlet boundary conditions at the AdS boundary leads to quantization: quasi-normal modes.

## Is anything wrong with viscous hydro?

The phenomenologically preferred value for  $\eta/s$  is very small  
Viscosity kills the elliptic flow!



P. Romatschke, U. Romatschke, Phys.Rev.Lett.99:172301,2007

Largely supported by H. Song, U. W Heinz, .arXiv:0712.3715

Somewhat disagrees with K. Dusling, D. Teaney arXiv:0710.5932

At RHIC Hydro seems to start at very early times  $\tau_0 \sim 0.5 - 1 \text{ fm}$ .

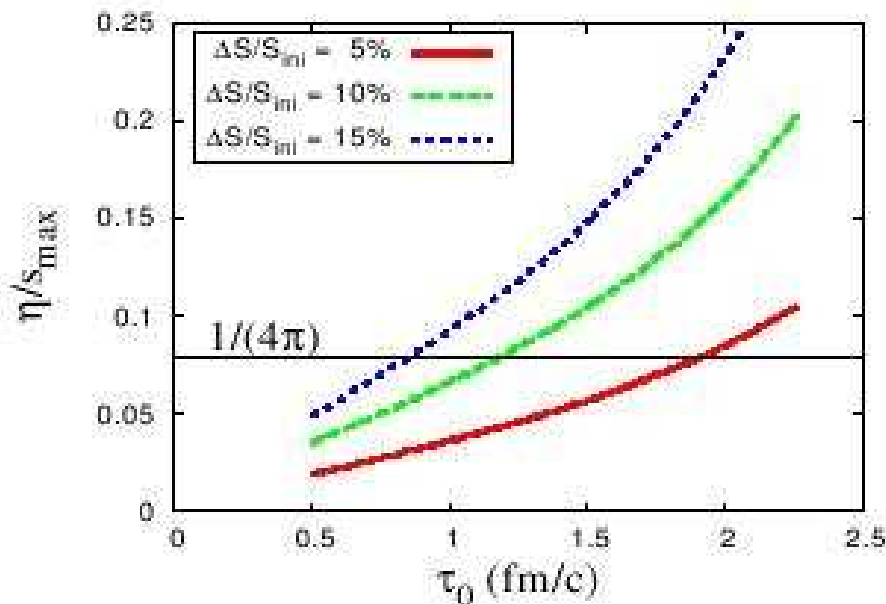
The hydro phase is long ( 10 fm)  $\rightarrow$  too much entropy is produced by the hydro phase.

That seems to contradict the RHIC data on produced particle multiplicities.

### Entropy production in the Bjorken (1d) Hydro

$$\partial_\tau (s \tau) = \frac{4 s}{3} \frac{\eta}{s} \frac{1}{T \tau}$$

A. Dumitru , E. Molnar , Y. Nara, Phys.Rev.C76:024910,2007.



E. Shuryak and M.L., Phys.Rev.C76:021901,2007

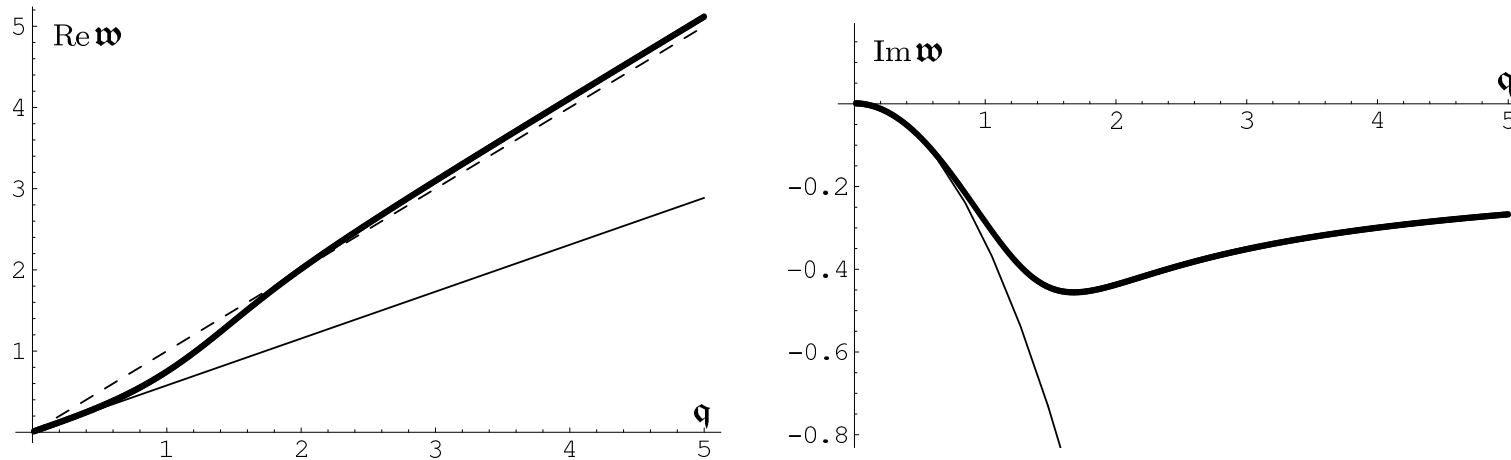
Introduce higher viscosity terms in  
the gradient expansion of  $T^{\mu\nu}$ :

$$\frac{\partial_\tau (s \tau)}{s (\tau T)} = \frac{\bar{\eta}}{\pi} \left[ c^2 \frac{1}{(\tau T)^2} + \sum_{n=2}^{\infty} \frac{\alpha_n}{(T \tau)^{2n}} \right]$$

# Sound and Holography

P. Kovtun and A. Starinets, Phys.Rev.D72:086009,2005

## Quasi-normal mode analysis in the AdS BH background - the sound channel



$$\Re[\omega] = c q + \sum_{n=1}^{\infty} r_n q^{2n+1}$$

$$\Im[\omega] = -\bar{\eta} \left[ c^2 q^2 + \sum_{n=2}^{\infty} \beta_n q^{2n} \right]$$

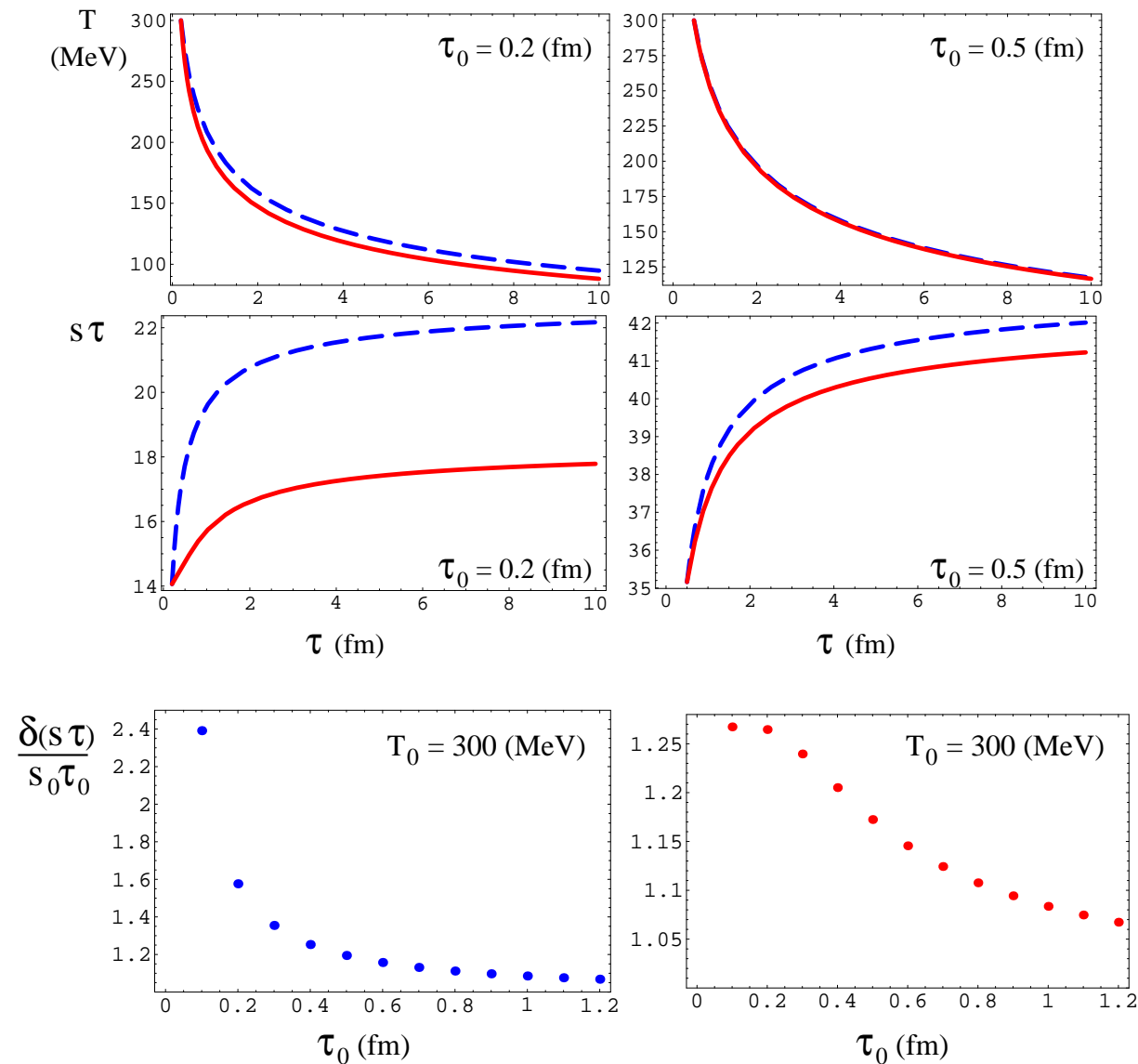
$$r_1 \rightarrow \tau_R = 2 - \ln[2]$$

R. Baier, P. Romatschke, D.T. Son, A.O. Starinets, M.A. Stephanov, arXiv:0712.2451

S. Bhattacharyya, V. E Hubeny, S. Minwalla, M. Rangamani, arXiv:0712.2456

$\beta_2 < 0$  while the IS second order hydro leads to  $\beta_2 > 0$

# How much entropy is produced by Hydro at RHIC?





## Linearized Hydro to all orders

$$\Pi^{\mu\nu} = \Delta^{\mu m} \Delta^{\nu n} D_{mn,k}[\nabla] u^k$$

**Tracelessness condition:**  $\Delta^{mn} D_{mn,k}[\nabla] u^k = 0$

$$D_{mn,k} u^k = g_{mn} \left[ \frac{2}{3} \eta_1 - \frac{1}{3} \eta_2 \nabla^2 \right] (\nabla u) - \eta_1 [g_{mk} \nabla_n + g_{nk} \nabla_m] u^k + \eta_2 \nabla_m \nabla_n (\nabla u)$$

$$\eta_{1,2} = \eta_{1,2}[(u\nabla), \nabla^2] \rightarrow \eta_{1,2}[i\omega, \omega^2 - q^2] = \Re \eta_{1,2} + \Im \eta_{1,2}$$

$$\eta_1[\omega \rightarrow 0, \mathbf{q} \rightarrow 0] \rightarrow \eta$$

**We keep the nonlinear dispersion to all orders, but**

**We neglect nonlinear interactions (though some terms could be recovered).**

**Shear (Diffusive) channel:**

$$G_R^T(\omega, q) = \frac{\eta_1 q^2/2}{-i\omega + \eta_1 q^2/2}$$

$$\chi^R = \Im m \ G_R^T$$

**Sound channel:**

$$G_R^L(\omega, q) = \frac{2i\omega c^2 q^2 \tilde{\eta} - c^2 q^2}{\omega^2 - c^2 q^2 + 2i\omega c^2 q^2 \tilde{\eta}}$$

$$\chi^L = \Im m \ G_R^L$$

$$\tilde{\eta} = \eta_1 + \eta_2 (\omega^2 - 2q^2)/4$$

**In order to extract  $\eta_{1,2}$  we have to invert this relations.**

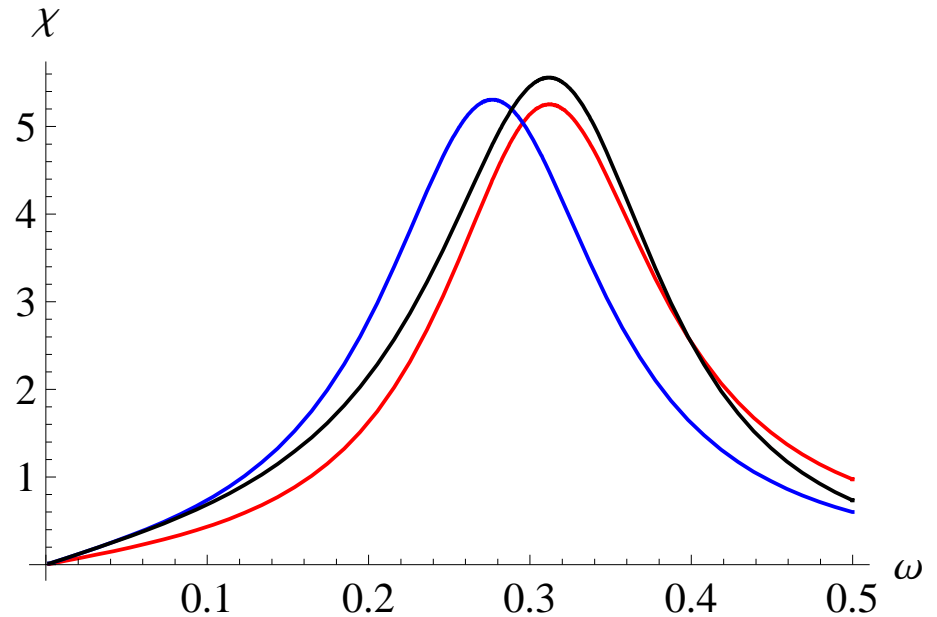
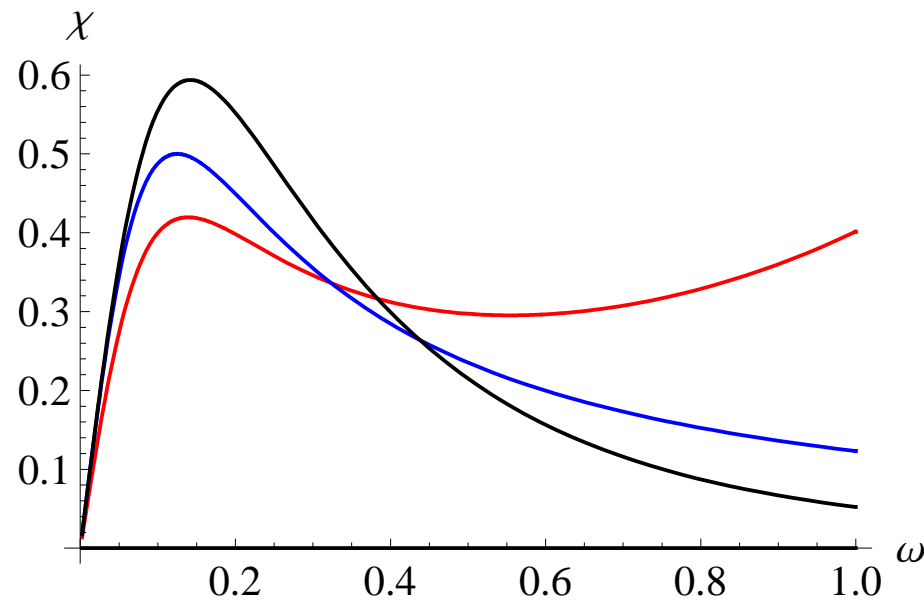
**For that we need both imaginary and real parts of the correlators.**

**Poles of the correlators should reproduce the entire tower of quasi-normal modes + their dispersion relations.**

Shear

$q = 0.5$

Sound



red - AdS BH

P. Kovtun, A. Starinets, Phys.Rev.Lett.96:131601,2006

blue - Navier Stokes

black - IS second order hydro.

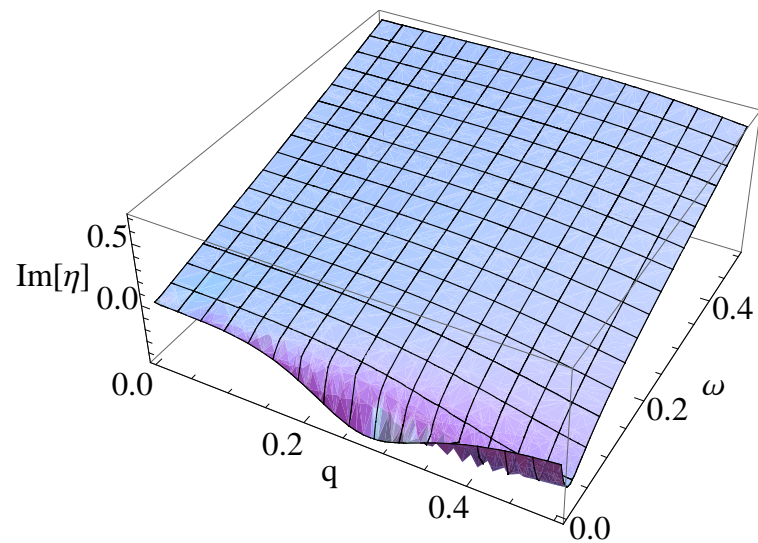
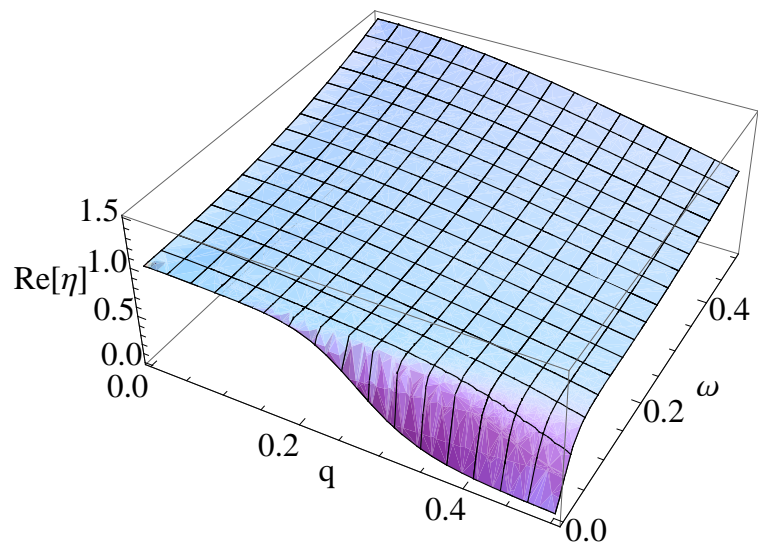
**Discussion point:**

The spectral functions contain “non-thermal” vacuum physics, such as pair production.

Should this physics be removed when constructing hydro?

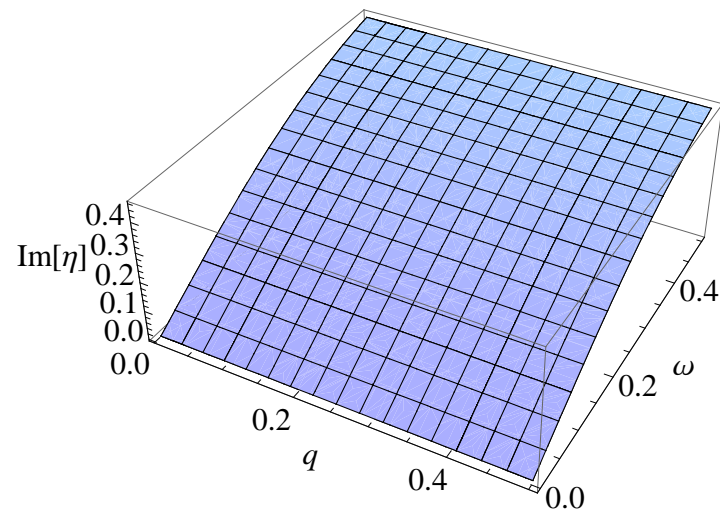
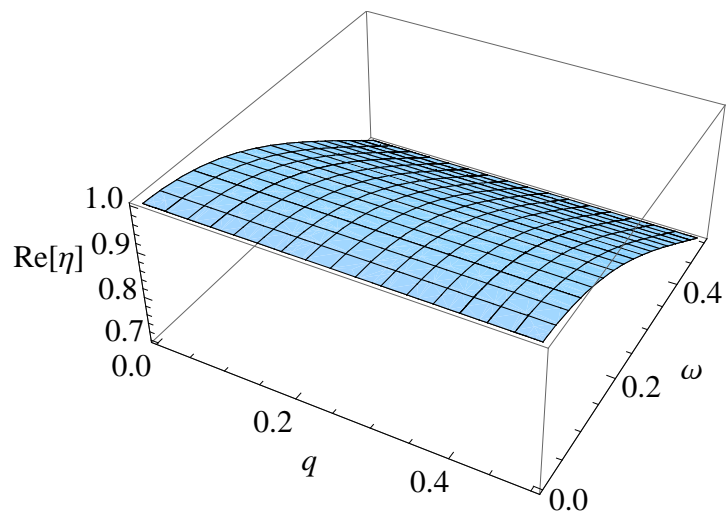
The “non-thermal” processes are real. They do occur in plasma.

Should we model them as an effective hydro?



$$\eta_1 = 1 + i \tau_R \omega + \kappa q^2 + \lambda \omega^2 \dots$$

$$\tau_R = 2 - \ln[2], \quad \kappa \simeq -1, \quad \lambda \simeq 1.7$$



$$\bar{\eta} = 1 + i \tau_R \omega - \tau_R^2 \omega^2 \dots$$

## Concluding Remarks

- Higher order terms in the gradient expansion seem to be important at early times. Taking them into account is likely to reduce the dependence on the initial time of the evolution.
- IS second order hydro does not agree with the all-order hydro from the AdS/CFT. This hints that this second order hydro is potentially less trustable tool than it could be previously thought.